Course Code: BSCP3001 Course Name: QUANTUM MECHANICS

Quantum Mechanics

Topic: Origin of Quantum Theory



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Covered Topics

- Black Body Radiation
- Photoelectric Effect
- Atomic Spectra
- ❖ Particle and Wave

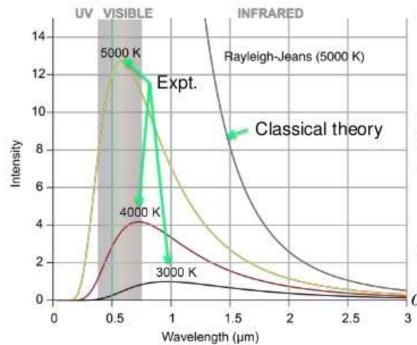
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Origin of Quantum Theory: Black Body Radiation

Black body radiation: when objects are heated, maximum wavelength of radiation shifts to shorter wavelengths.



Classical: Rayleigh-Jeans law

$$d\rho(\nu, T) = \frac{8\pi k_{\rm B} T}{c^3} \nu^2 d\nu$$

Planck (1900): Energies of oscillations in the black body are discrete or quantized!

$$E = nh\nu$$

Planck's constant

Planck distribution law:

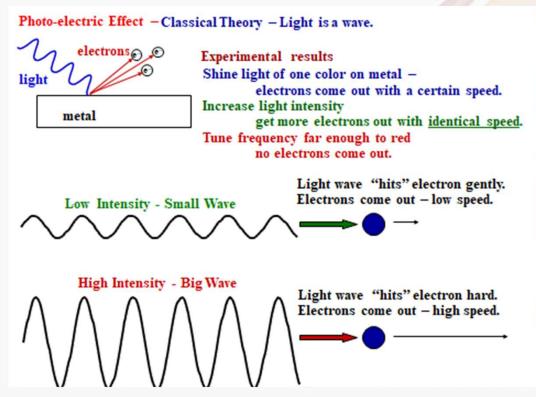
$$d\rho(\nu,T) = \frac{8\pi h}{c^3} \frac{\nu^3 d\nu}{e^{h\nu/k_B T} - 1}$$

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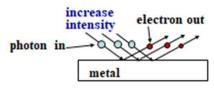
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Photoelectric Effect- The Nature of Light



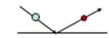
Einstein explains the photoelectric effect (1905) Light is composed of small particles – photons.



One photon hits one electron.

Increase intensity – more photons, more electrons hit – more come out.

Each photon hits an electron with same impact whether there are many or few.

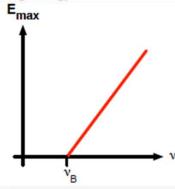


Therefore, electrons come out with same speed independent of the intensity.

Tune to red, energy to low to overcome binding energy.

$$K = h\nu - W = h(\nu - \nu_0),$$

$$V_s = \frac{h}{e}v - \frac{W}{e} = \frac{hc}{e\lambda} - \frac{W}{e}.$$

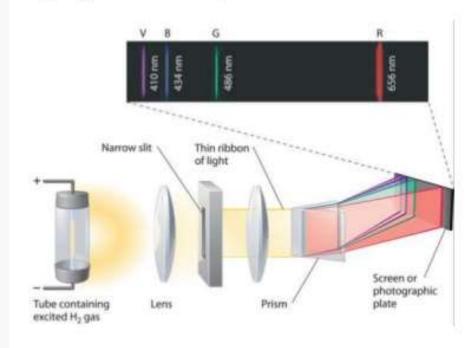


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Spectrum – Hydrogen Atom

Hydrogen atom line spectrum:



Rydberg (1888) discrete lines of hydrogen spectrum fit formula:

$$\frac{1}{\lambda} = R_{\rm H} \left(\frac{1}{n_1^2} - \frac{1}{n_2^2} \right)$$

where:

$$n_2 > n_1 = 1, 2, 3, 4...$$

Bohr (1911) model:

- 1) stationary electron orbit,
- 2) integral # of wavelengths

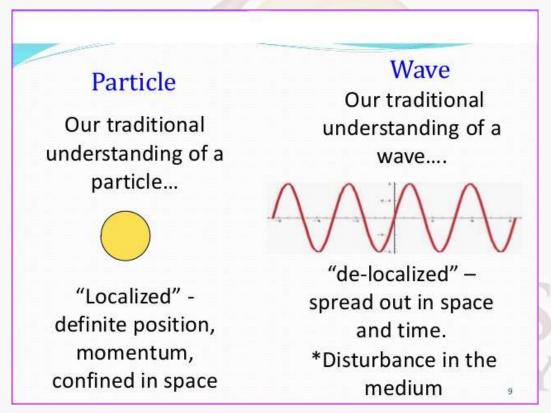
$$E_n = -\frac{m_e e^4}{8\varepsilon_0^2 h^2} \frac{1}{n^2}$$

where:

$$n = 1, 2, \dots$$

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Particle and Wave



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